

ATHENA

The Athena mission

Status and Community Preparation

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Thanks

- Special thanks:

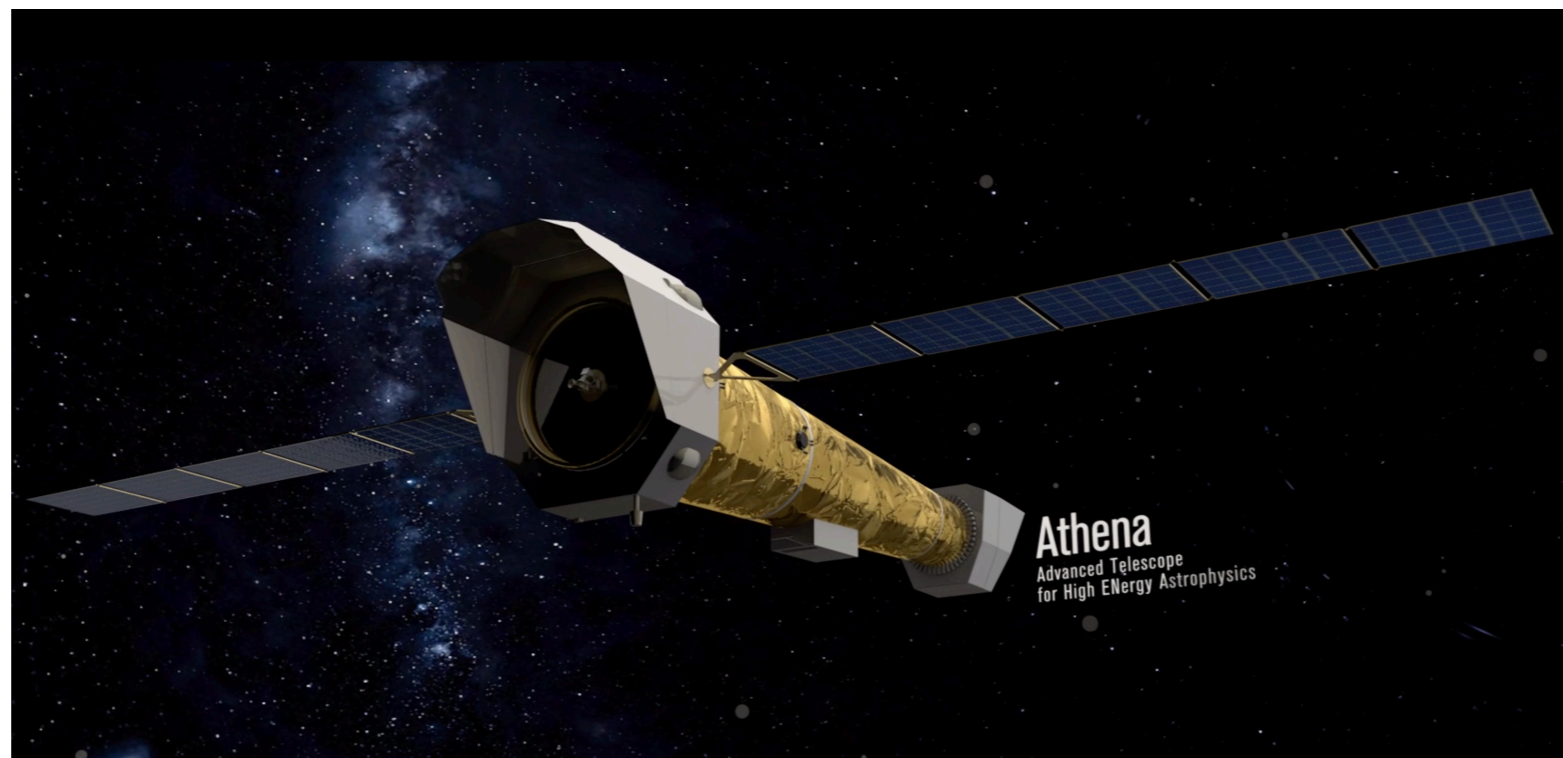
- ▶ to the **Athena Science Study Team**: D. Barret, A. Decourchelle, A.C. Fabian, **M. Guainazzi**, J.W. den Herder, H. Matsumoto, K. Nandra, L. Piro, R. Smith, R. Willingale
- ▶ to the Athena Working Group and Topical Panel chairs and members
- ▶ to the ESA study team
- ▶ to the **Athena Community Office (ACO)**
- ▶ to the WFI instrument consortium
- ▶ to the **X-ray Integral Field Unit Consortium**

Outline

- Brief recap of what Athena is
- Project status, including X-IFU and technology development status
- Athena organization and community involvement, including some aspects of the Athena science management plan
- Conclusions: **We are on safe track for a launch in early 2030s**

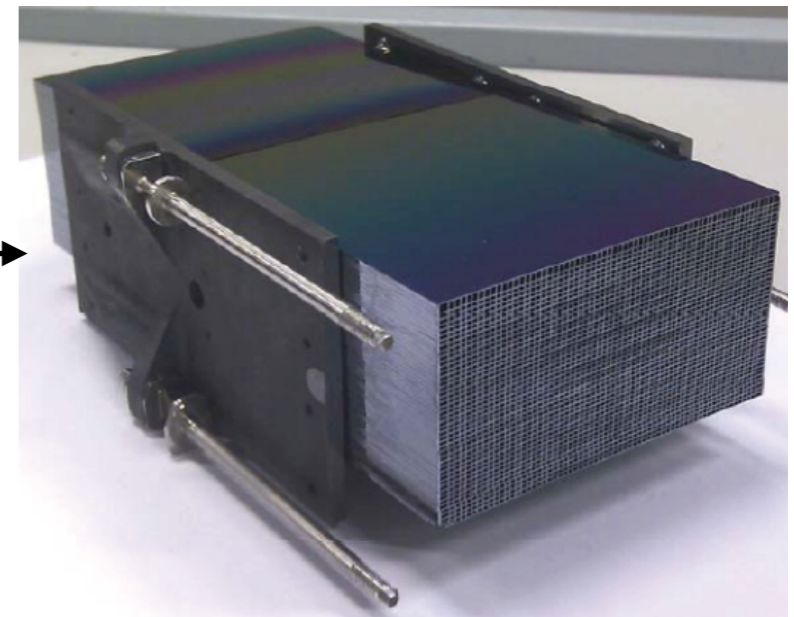
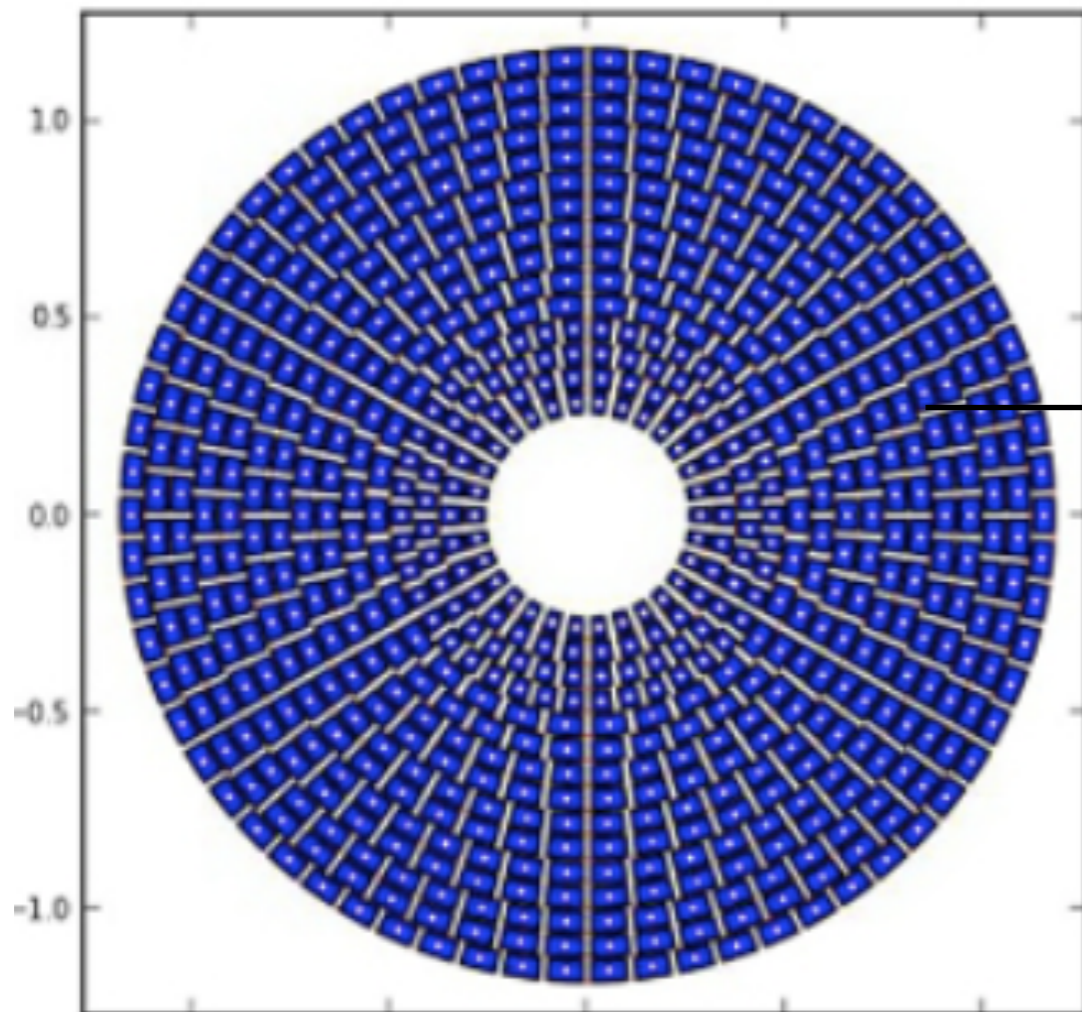
Athena

- Athena: **A**dvanced **T**elescope for **H**igh **E**Nergy **A**strophysics
- **S**econd **L**arge **m**ission of the European Space Agency Cosmic Vision Science program (the third being the LISA gravitational wave mission)
- Dedicated to **T**he **H**ot and **E**nergetic **U**niverse
 - ▶ With broad impacts in many corners of astrophysics: stars, galaxies, planets... which define the **O**bservatory **s**cience of Athena
- One large aperture telescope and two complementary instruments



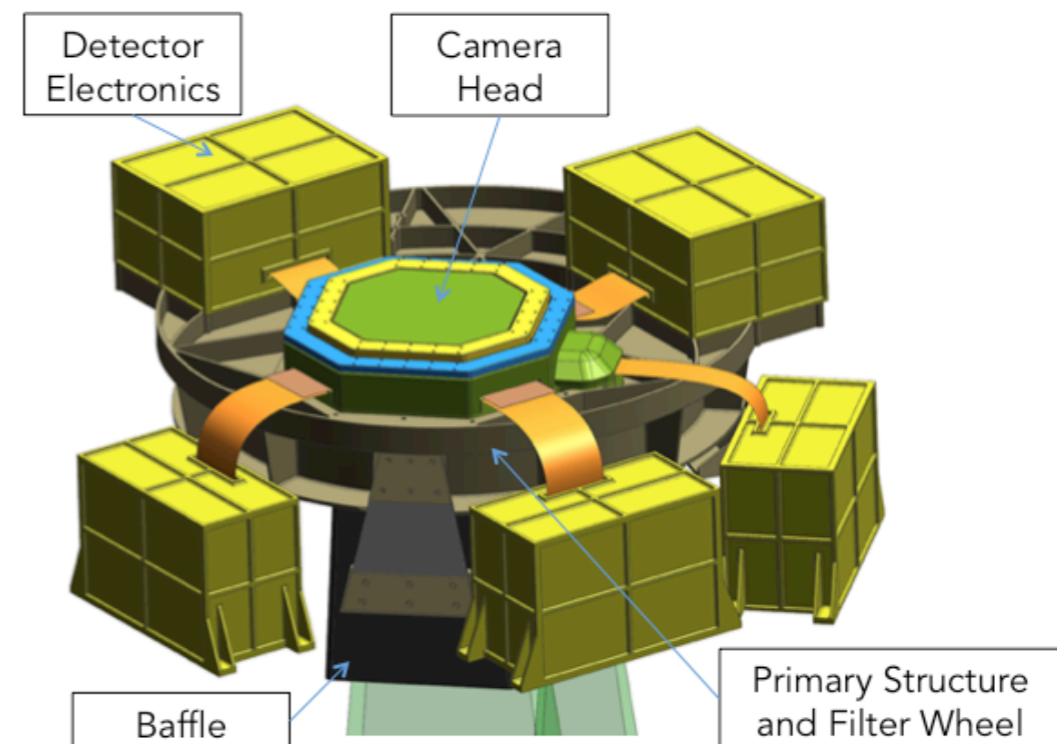
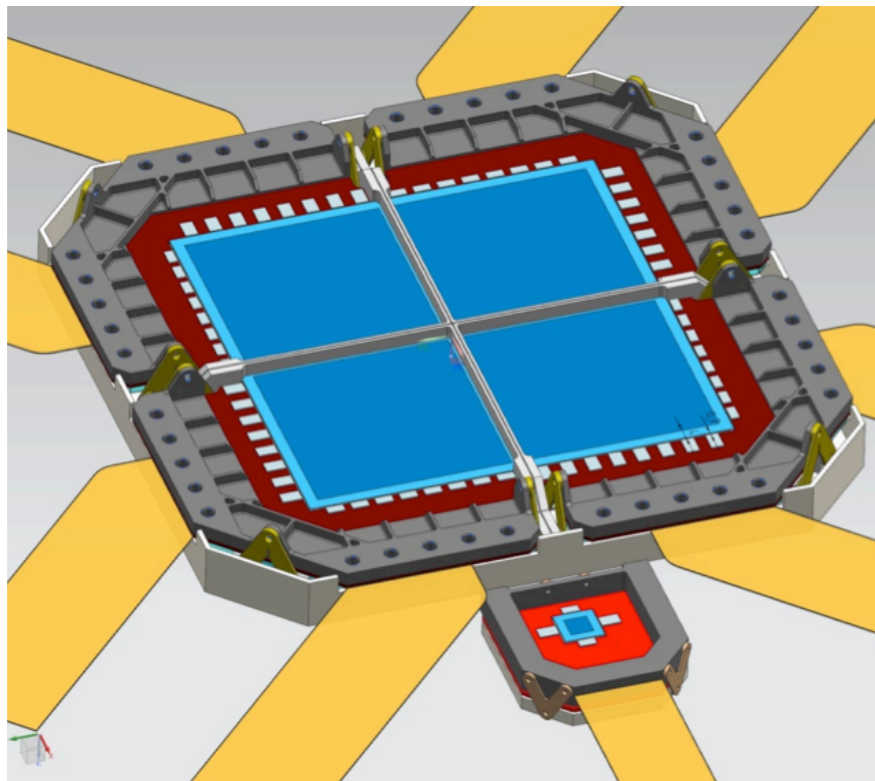
The Athena science payload : Mirror

- Single monolithic large aperture grazing incidence **movable** X-ray telescope
 - ▶ **Silicon Pore Optics** developed by ESA
 - ▶ **1.4 m² @ 1 keV, 0.25 m² @ 6 keV** (requirements)
 - ▶ **5'' (HEW)** requirement



The Athena science payload: WFI

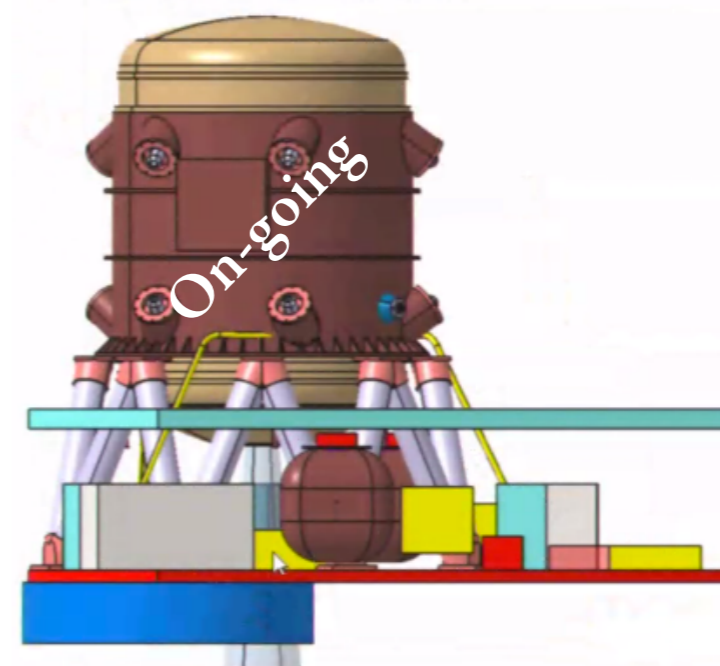
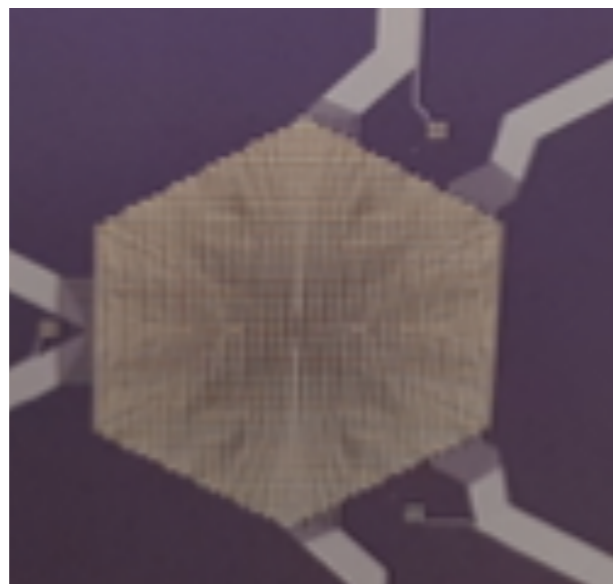
- **Wide Field Imager (WFI) — PI K. Nandra (MPE)**
 - ▶ Silicon Active Pixel Detector based on DEPFET technology
 - ▶ <80 (<170) eV spectral resolution @ 1 (7) keV
 - ▶ 2.2'' pixel size (PSF oversampling)
 - ▶ Field of view: 40' \times 40' square
 - ▶ Separate chip for fast readout of brightest sources
 - ▶ Consortium led by MPE, with other European partners (DE, AT, DK, FR, IT, PL, UK, CH, P & GR) and NASA



Credits: MPE and WFI team

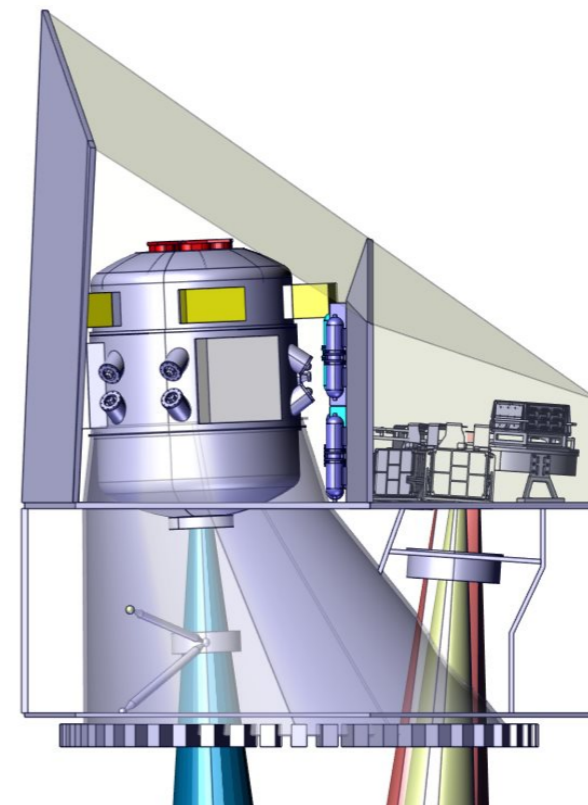
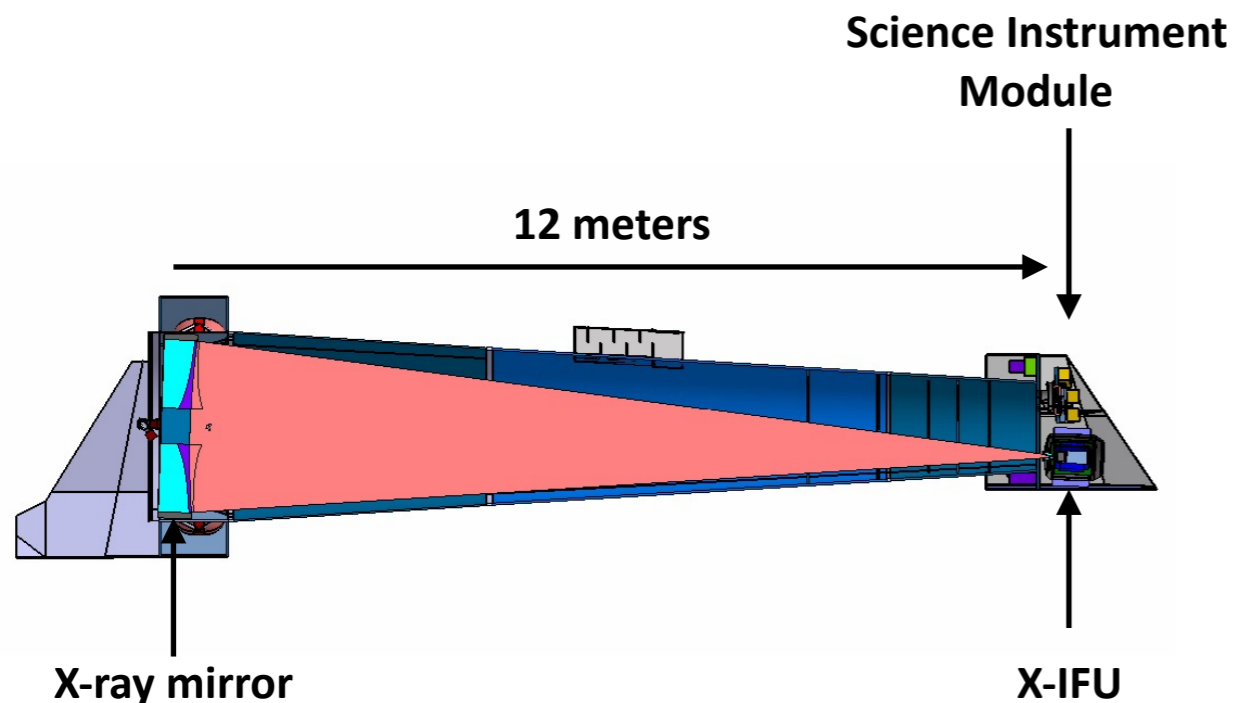
The Athena science payload : X-IFU

- **X-ray Integral Field Unit (X-IFU):** Co-PIs: J.W den Herder (SRON) & L. Piro (INAF)
 - ▶ Large format micro-calorimeter array (Transition Edge Sensors)
 - ▶ **2.5 eV** spectral resolution up to 7 keV (frequency domain multiplexing)
 - ▶ **5'** hexagonal field of view (equivalent diameter)
 - ▶ Low background due to a cryogenic anti-coincidence detector
 - ▶ Capability to observe bright sources (1 Crab) thanks to the mirror defocussing
 - ▶ **Cryogenic instrument** cooled down to 50 mK by a multi-stage cryogenic chain
 - ▶ Consortium led by IRAP/CNES-F, with NL and IT and further ESA member state contributions from BE, CZ, FI, DE, IR, PL, ES, CH and contributions from Japan and the United States

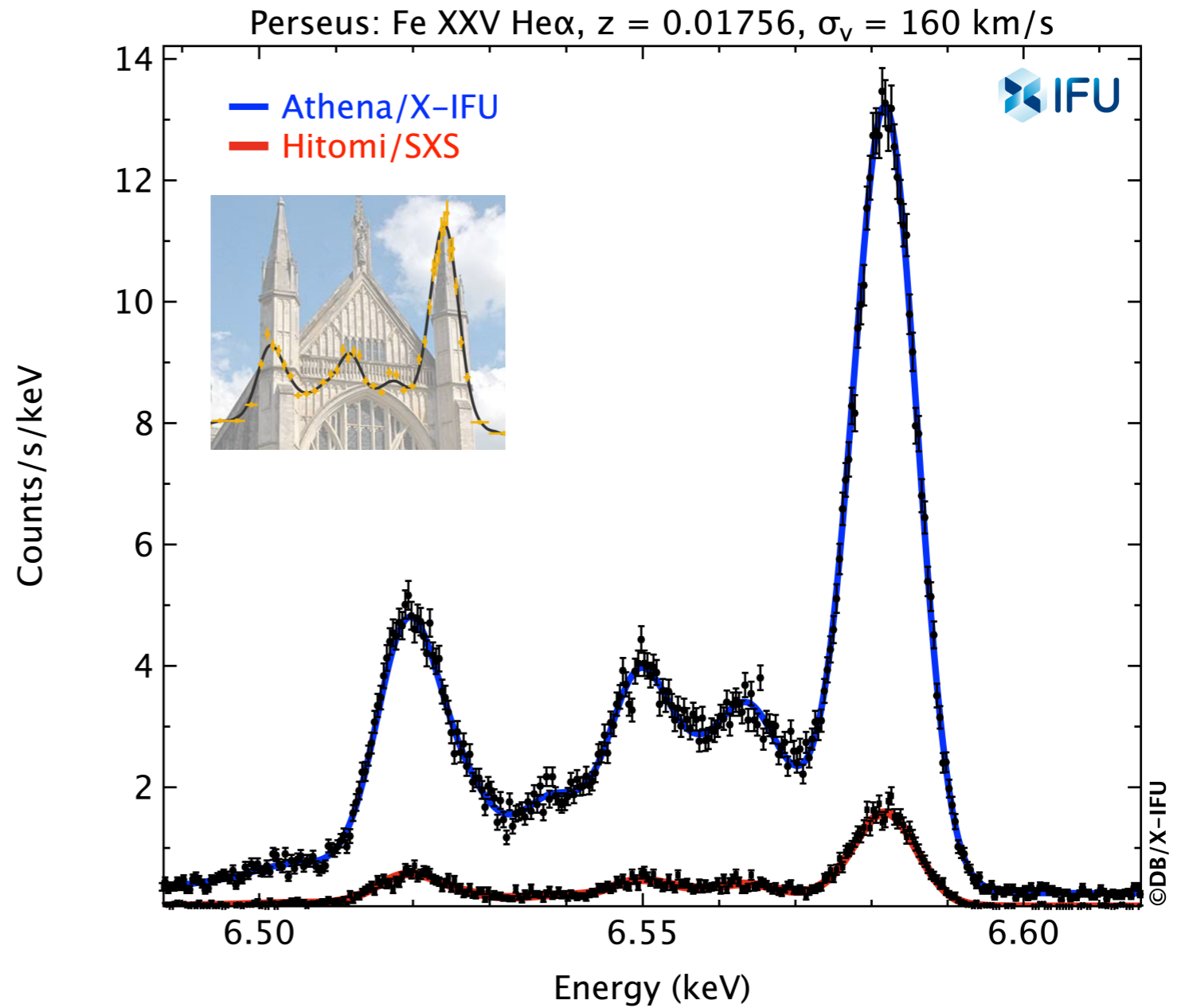
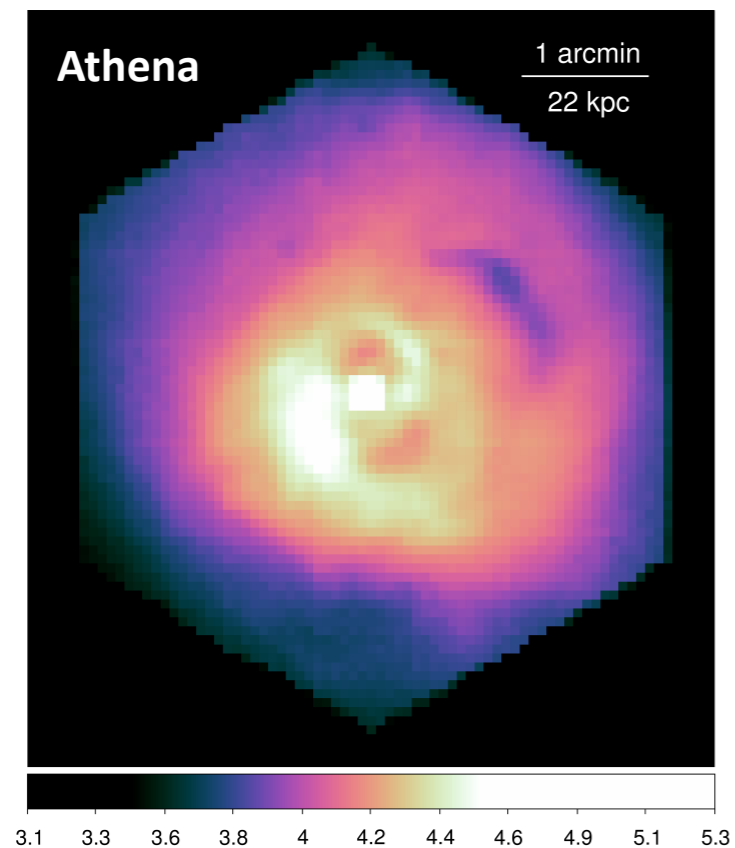
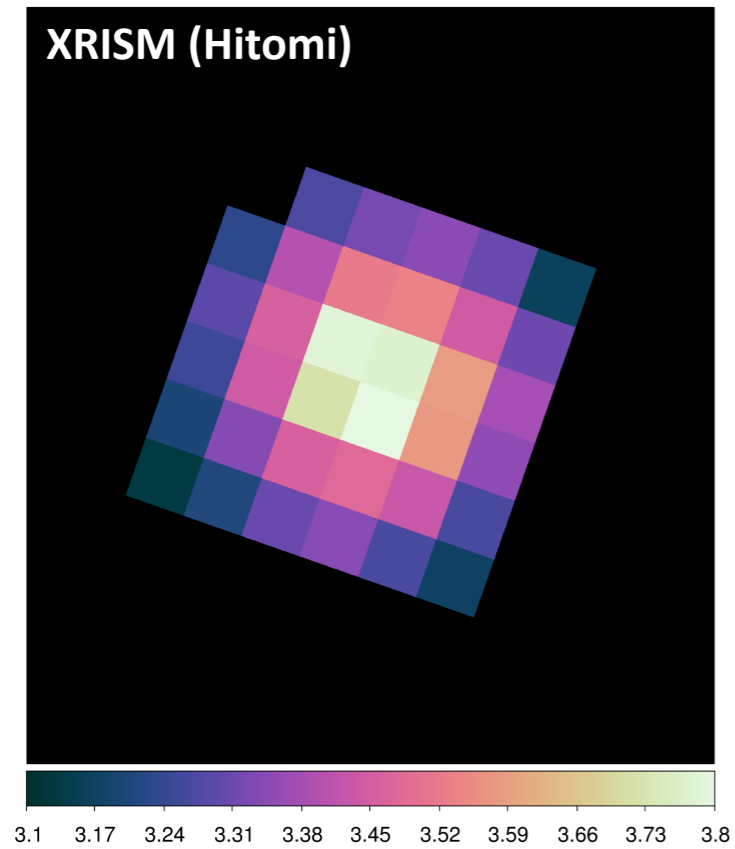


The Athena spacecraft

- Focal length: 12 meters (total height about 15 meters)
- Overall mass: 7 tons (X-IFU ~ 1 ton)
- 7 kWatts
- 4 year nominal mission lifetime with **consumables/mechanical parts** designed for **10 years**
- Agile satellite to respond to ToO alerts in a few hours
- Launch to halo orbit L2 (or L1) by Ariane 6



From XRISM to Athena: Perseus



Credits: J. Sanders

Project status

- Instrument consortia formally appointed by ESA (Dec 2018)
- Both instruments **successfully** completed their feasibility study phase (Phase A) and can now proceed to phase B1 (since April 2019)
 - ▶ Their baseline designs **meet** the performance requirements
 - ▶ Actions raised by the review are being implemented with the goal of:
 - ▶ consolidating our baseline designs and interfaces with the spacecraft
 - ▶ consolidating and running our technology developments plans
 - ▶ consolidating the programmatic aspects
- Next review to come (Sept. 2019): End of phase A review at mission level (including mirror and spacecraft as ESA led elements), so-called Mission Formulation Review (MFR)
 - ▶ According to the ESA study team: *Athena is in good shape for MFR*

X-IFU status

- X-IFU passed the IPRR with a **baseline** declared **feasible**
 - ▶ Micro-vibrations and cooling chain complexity identified as risks
- Delta-study conducted by ESA & CNES to see how the science instrument module could ease the accommodation of the X-IFU
 - ▶ By providing more passive cooling to the instrument in a so-called « cryo-SIM » configuration
- Management meeting of July 8th concluded that the baseline **configuration is confirmed** weighting benefits versus overall system complexity
- No disruption of activities, and optimizations are now under investigation as part of the normal phase B1
 - ▶ To address micro-vibrations
 - ▶ To reduce the number of coolers
 - ▶ **At iso-performances**

Technology: optics

- Vigorous optics development plan on-going at ESA/Cosine:
 - ▶ Steady and continuous improvement in the optics towards 5" HEW
- A sensitivity analysis led by the ASST concluded that a resolution of 6.5" has a clear negative impact on a number of Athena science objectives, with the impact considered very severe if the HEW were to degrade to >8"
 - ▶ WFI science more severely impacted than X-IFU science
 - ▶ Today 7" is achieved on 70% of the plate area
- Angular resolution optimization activities will continue beyond MFR
- Various coatings for the optics under investigation (Ir, Ir+SiC), as well as ways of maximizing the effective area of the mirror at both 1 and 7 keV

Technology: instruments

- Prime focus is X-IFU
- Consistent assessment of the level of readiness of X-IFU critical technologies between ESA and X-IFU team
 - ▶ On track for mission adoption although schedule is tight
 - ▶ Significant progresses made towards reaching 2.5 eV resolution with the baseline frequency domain multiplexing
 - ▶ System level demonstration to be performed by the demonstrator of the X-IFU cooling chain to include a demonstrator of the focal plane assembly (32 x 32 TES array) and readout electronics
- Technology demonstration to be completed by mission adoption in Q3-4/21
- 10 year implementation time to launch in early 2030's
 - ▶ 6 month provision of the calibration time for X-IFU !

Athena organization

- The Athena Science Study Team is an ESA-appointed body of astronomers, including the ESA Study Scientist, the instrument PIs and international representatives from JAXA and NASA.
 - ▶ Define the scientific requirements of the mission and to advise ESA on all aspects related to the scientific performance of Athena.
- The ASST has nominated 5 Working Groups on:
 - ▶ Hot Universe, Energetic Universe, Observatory Science, Telescope, Mission performance
 - ▶ And 21 topical panels covering sub-topics
 - ▶ **Yearly calls enable you to join**
- The WG advise the ASST which in turn advises the ESA study team
 - ▶ e.g. consolidating science requirements, evaluating the impact of performance changes, providing inputs on calibration, etc.
 - ▶ Refinements to the organization to be discussed with the Chairs (Dec 2019)

Athena Community Office (ACO)

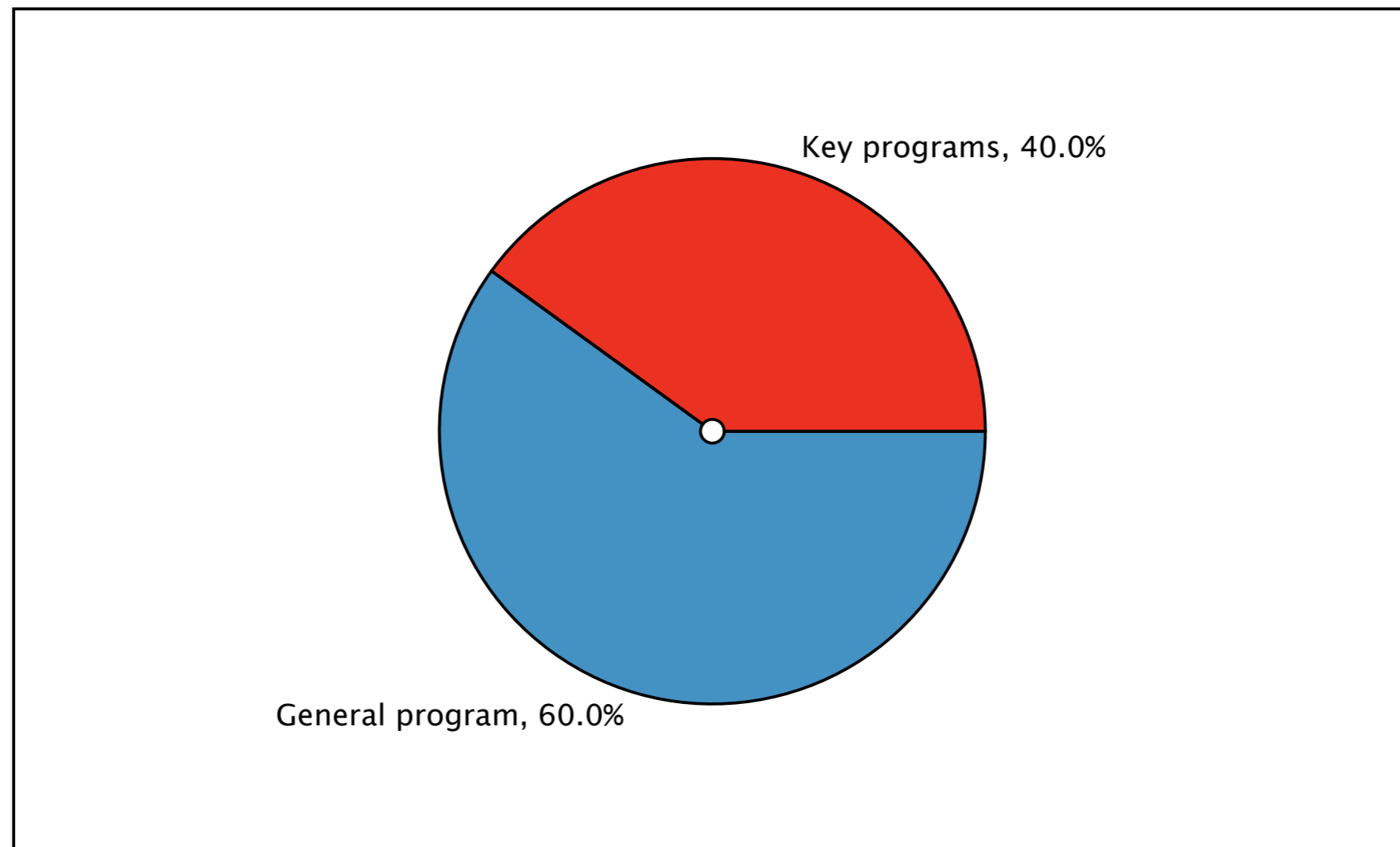
- The Athena community gathers today about 800 researchers
 - ▶ Bringing their expertise and the advises to the project is key
 - ▶ Major effort required to get the community involved
- To run it, the ASST appointed the Athena Community Office to :
 - ▶ Organize and optimize community efforts
 - ▶ Pass on the information from the ASST to the community
 - ▶ Develop communication and outreach activities around Athena
 - ▶ Led by IFCA (CSIC-UC) in Spain, with contributions from IRAP, MPE and UniGe
- More info at www.the-athena-x-ray-observatory.eu and on the social networks

How to get engaged and prepared?

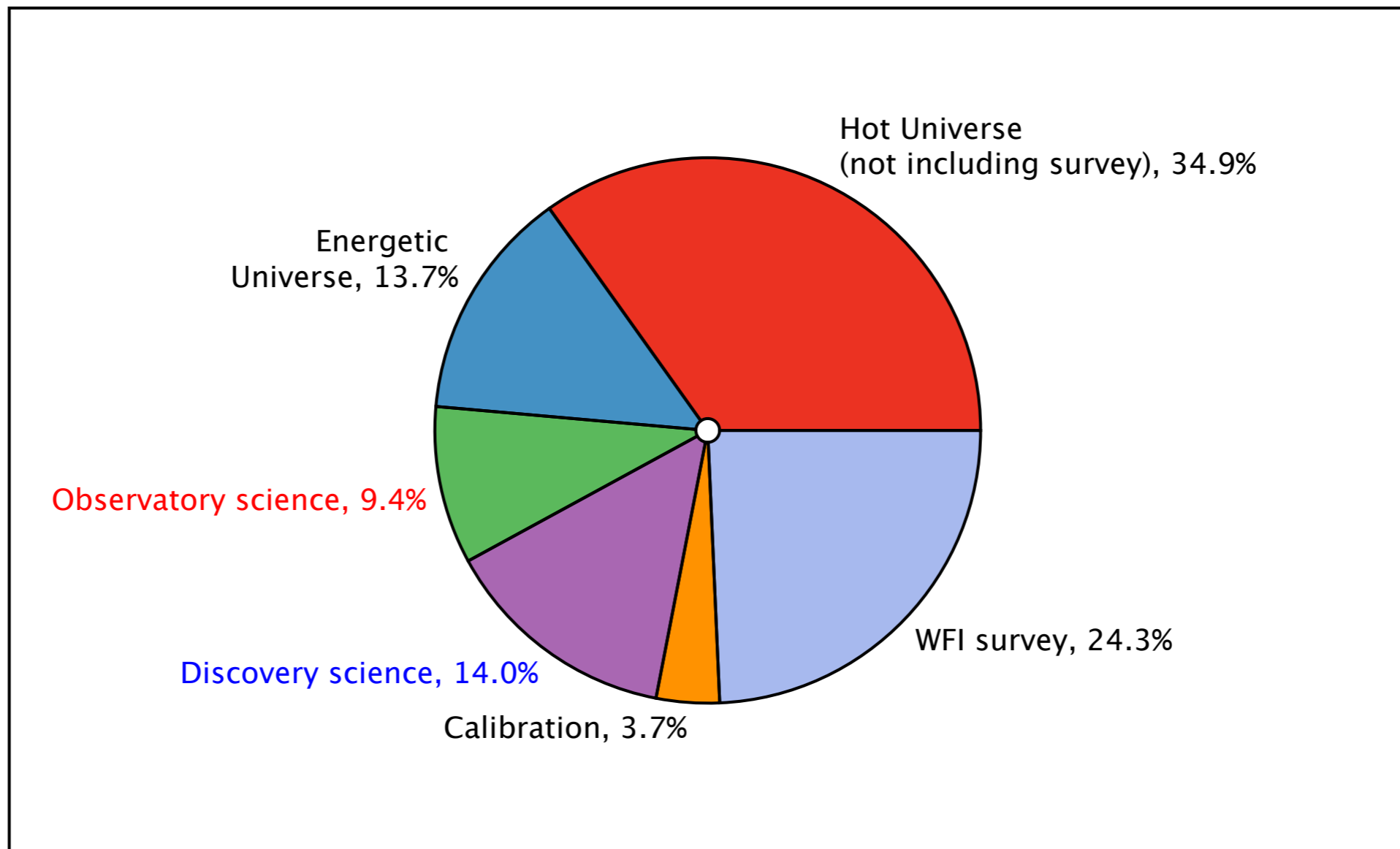
- Joining the consortia providing the instruments and their associated Instrument Science Centers (ISCs);
 - ▶ A bit late for the instruments, but opportunities may remain for the ISCs
- Joining or being part of one of the topical panels
 - ▶ Getting trained through SIXTE workshop
 - ▶ Participating to activities: e.g. a series of refereed journal papers will be prepared in time for mission adoption (for the so-called red book):
 - ▶ List of papers will be based on the received topical panel chair proposals
 - ▶ Complete and uniform coverage of existing science goals of Athena, as well as possible new science ideas discussed
- Get to work or /support XRISM/Resolve data, ensuring that special needs for X-IFU will be properly fed to the system (e.g. new tools)
- Becoming an Athena guest observers, and joining Key Programme team and in their preparatory work

Type of observing programs

- KP require the study of well-chosen samples of objects with particular properties
- Key Programs (KP) will be defined before launch by an ad'hoc committee taking consideration of the evolution of the scientific landscape

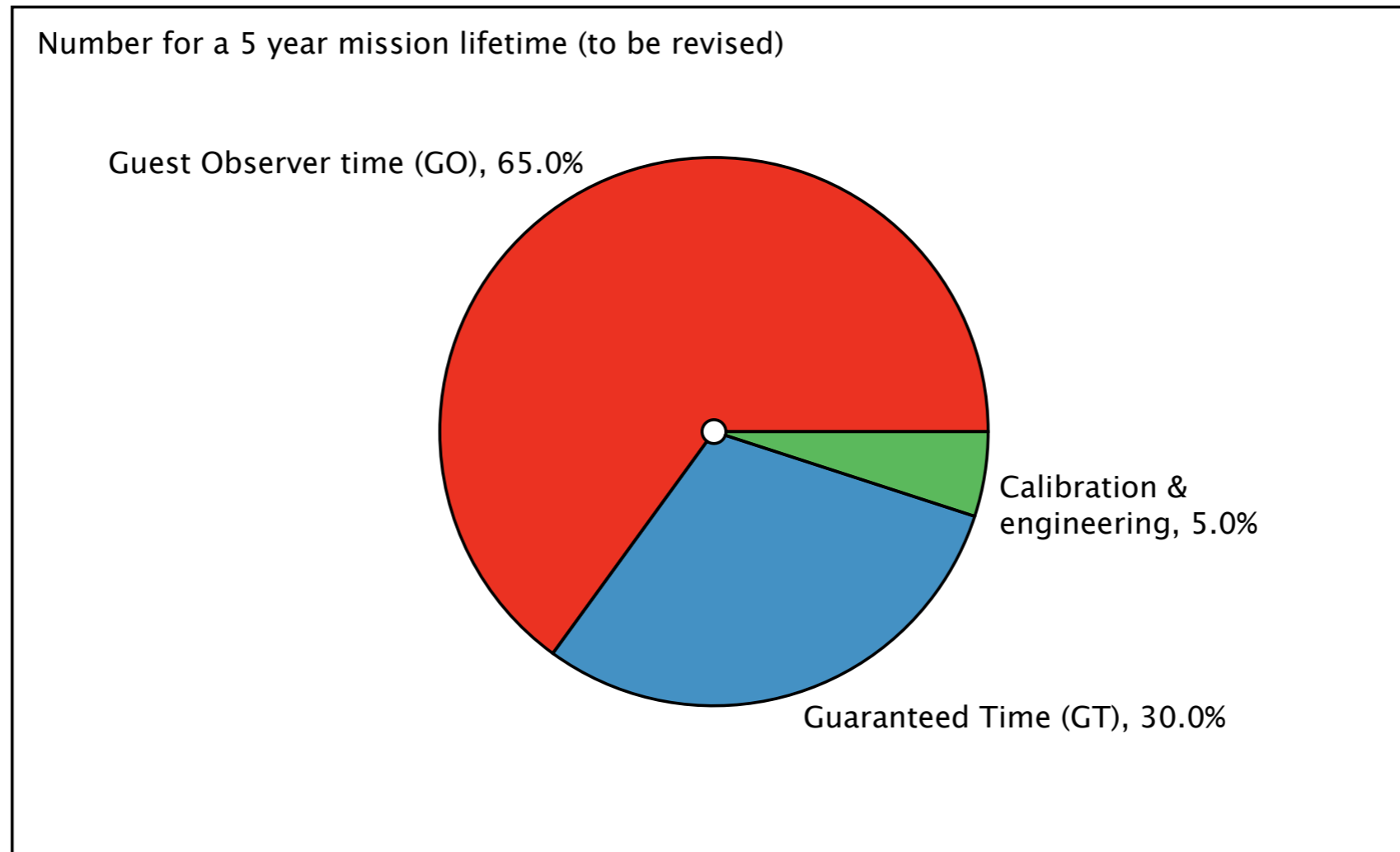


Observing time split



- Core science program defines **today** about 73 % of the available observing time (=107 Ms over a 4 year mission lifetime with 85% efficiency)
 - ▶ Numbers to be taken as indicative as they depend on the mission performance

Types of observing times



- Guaranteed Time (24%, tbc) and Guest Observer time (16%) to be spent on Key Programs, the rest being used for the General Program
 - ▶ GT will reduce to 5% (tbc) during mission extension
- Pending **approval by the ESA Science Program Committee** at the time of mission adoption (Q4/2021)

Conclusions

- Athena is **your** next large X-ray observatory
- Athena has **revolutionary** X-ray capabilities
 - ▶ Spatially resolved high resolution spectroscopy
 - ▶ Wide field imaging
- **XRISM** will open a new era of X-ray astronomy
 - ▶ Challenges for XRISM/Athena: calibration, background (+CX), atomic physics, data analysis
- Athena is now on the **safe path** for a launch in the early 2030s
 - ▶ Case is made to have **overlap** between Athena and **LISA**
- Keep using XMM-Newton/Chandra, get ready for XRISM, and keep **supporting and promoting** Athena
 - ▶ And start now preparing for the post-Athena mission in the **voyage2050** for the long term science planning of ESA